

# Neutron star evolutions using tabulated equations of state with a new execution model

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# Motivation

- Adding nuclear and neutrino physics to GR fluid codes gives a more realistic description of hot nuclear matter. Asynchronous nonblocking access to large tables of data is needed.
- Transit the pan-Petaflops performance regime to sustained Exaflops before the end of this decade.
- Improve strong scaling in scaling constrained codes by employing a modern execution model.
- Address issues of programmability in order to realize practical Exascale processing capability



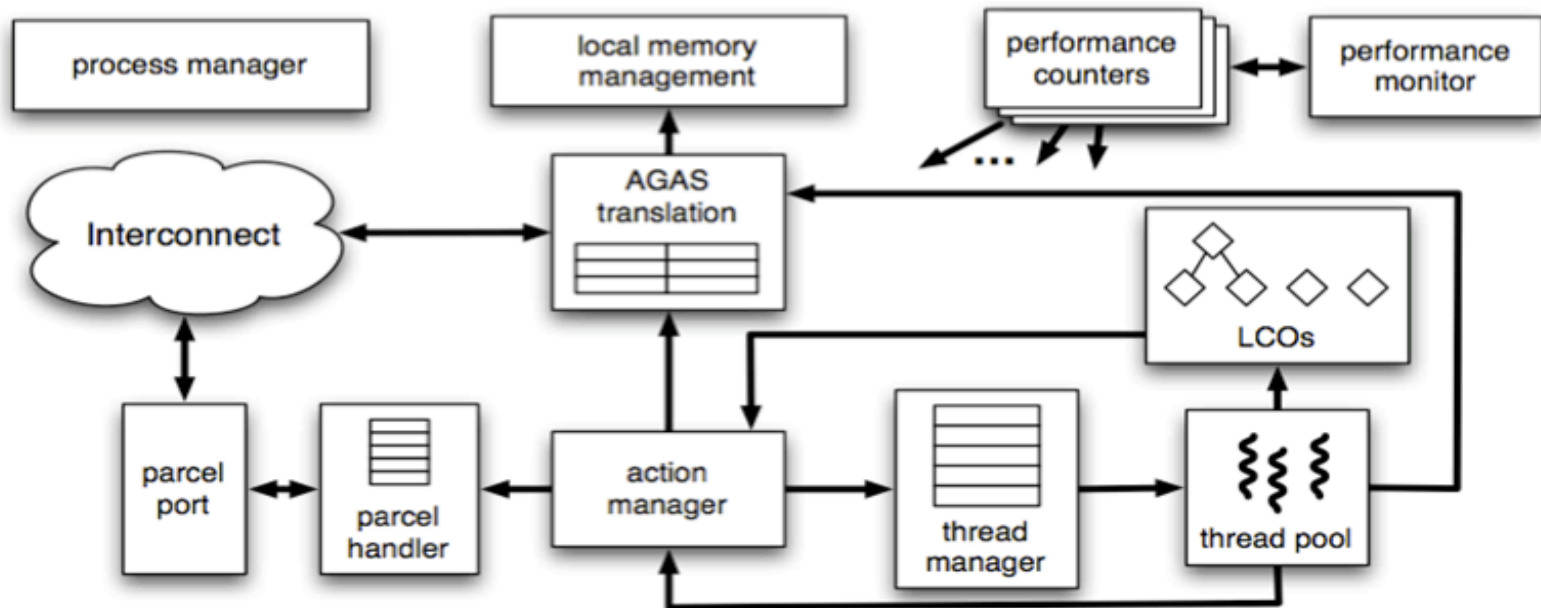
# Overview

- A little about ParalleX
- Performance comparisons using ParalleX for the EOS routines and tables found at <http://stellarcollapse.org/equationofstate> (O'Connor and Ott)
- Towards Radial pulsation frequencies



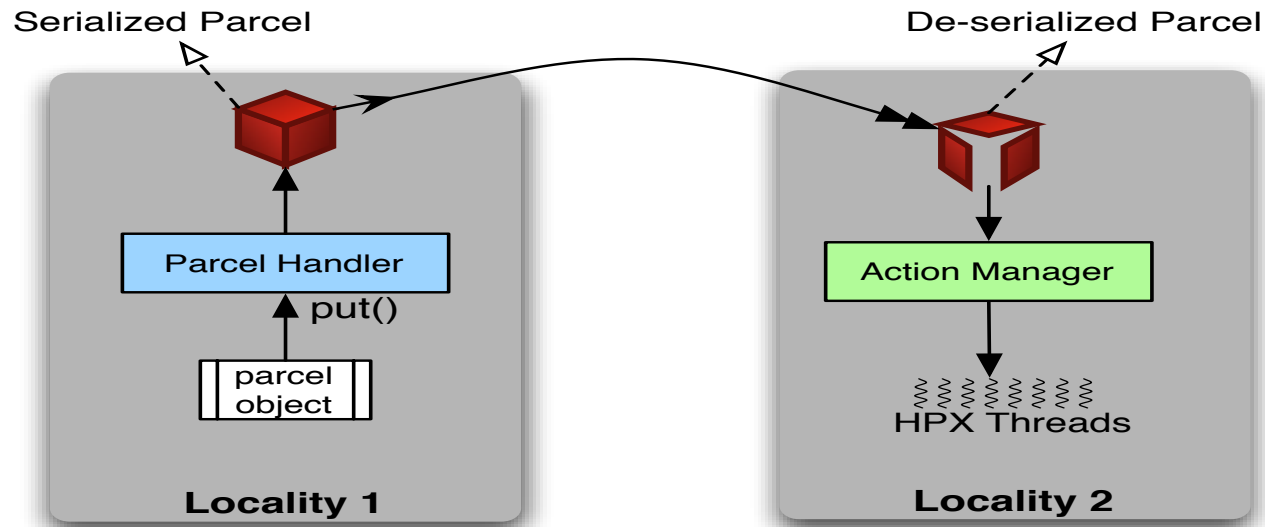
# ParallelX

- Establishes a global address space that is active in the sense that a virtually addressed object may migrate across nodes without having to change its address



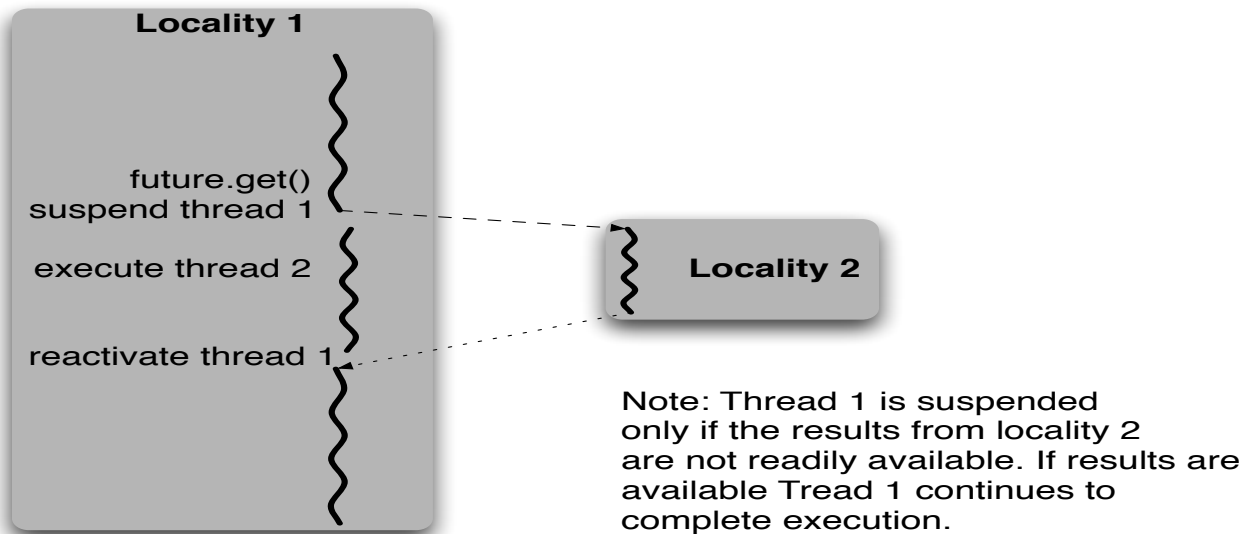
# ParalleX

- Communicates by parcels (an advanced form of active messages) that moves both work and control state to the data as well as conventional asynchronous gathers of data to the work



# ParalleX

- Supports Local Control Objects (LCOs) for lightweight synchronization to eliminate global barriers and manage asynchronous compound actions to reduce overhead and expose additional parallelism



# ParalleX

- Defines contexts of data and tasks that provide protected abstract domains across multiple system nodes.
- Thread manager implements work stealing (compare to Cilk)
- Open source implementation: HPX
- <http://stellar.cct.lsu.edu/>
- Shen EOS component used for this talk is available for download
- AMR code with tapered boundaries also available



# Finite Temperature Equation of State

- Putting in the right nuclear physics (polytropes don't have the right compactness).
- To do anything with neutrinos, you need a temperature: neutrino cooling, dynamics of hypernova
- To do anything with radiation, you need a physical temperature
- This is a spring-board for new astrophysics and microphysics



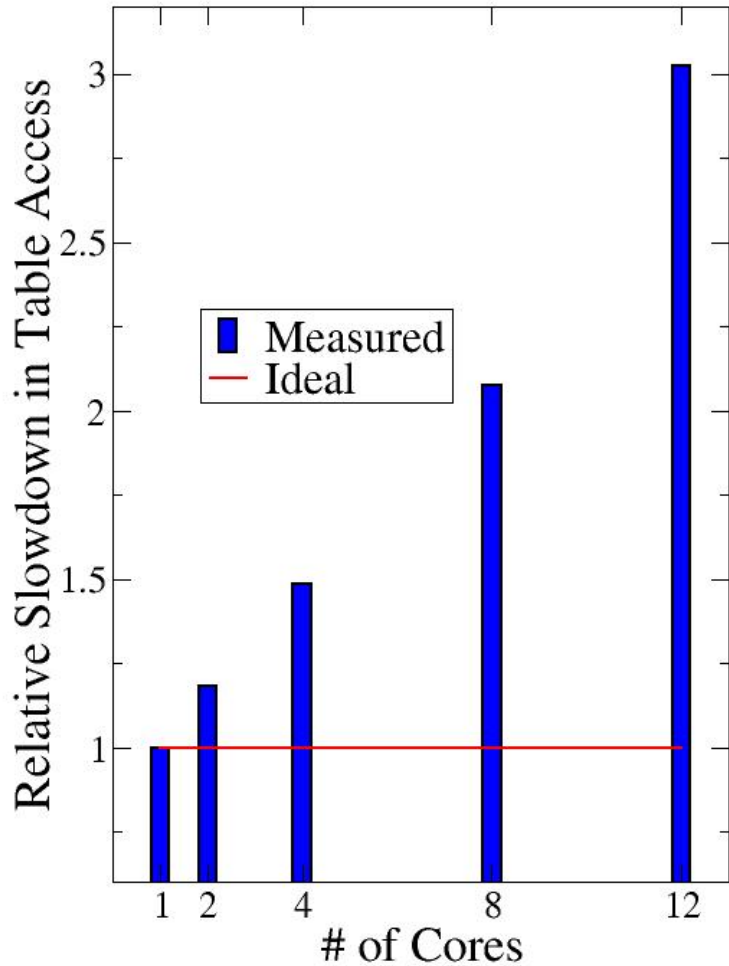


# Equations of State come in tables

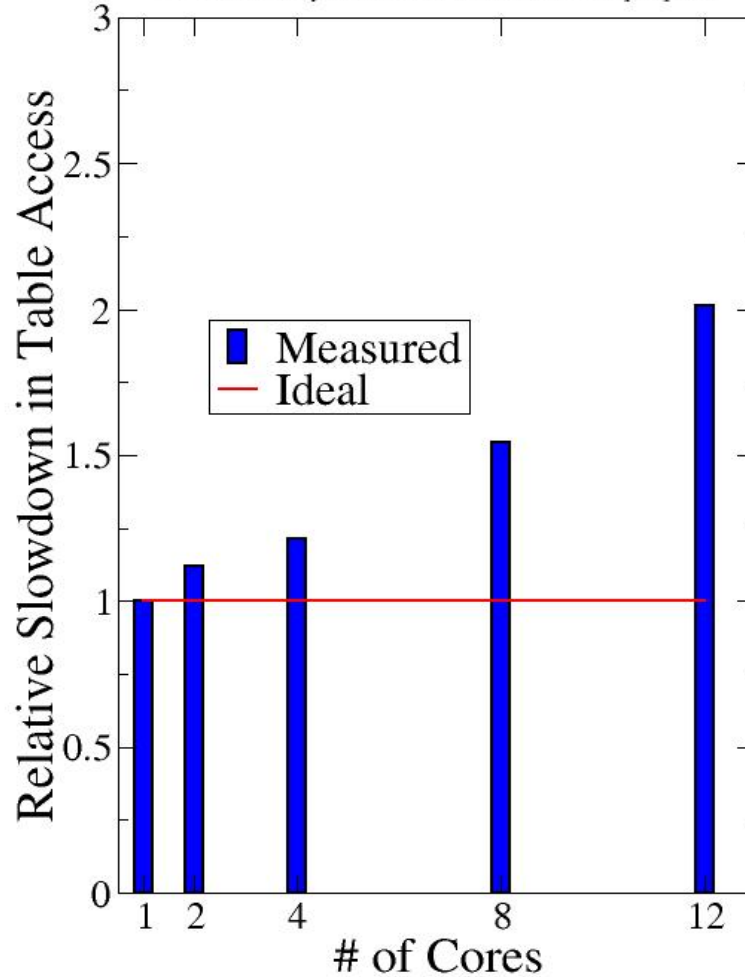
- Not practical to do EOS calculation in place; it's best to use a look-up table
- The table covers a lot of physics for you
- In high energy astrophysics:
  - Temperatures vary from 0 to  $> 100$  MeV
  - Proton fraction changes from 0 to 0.6
  - Density varies across 10 orders of magnitude
- Finer grid tables are better for accuracy
- Tables need to cover a wide variety of conditions: black hole formation, neutron star mergers, nucleosynthesis



Shen EOS Tabular Access  
Shared Memory

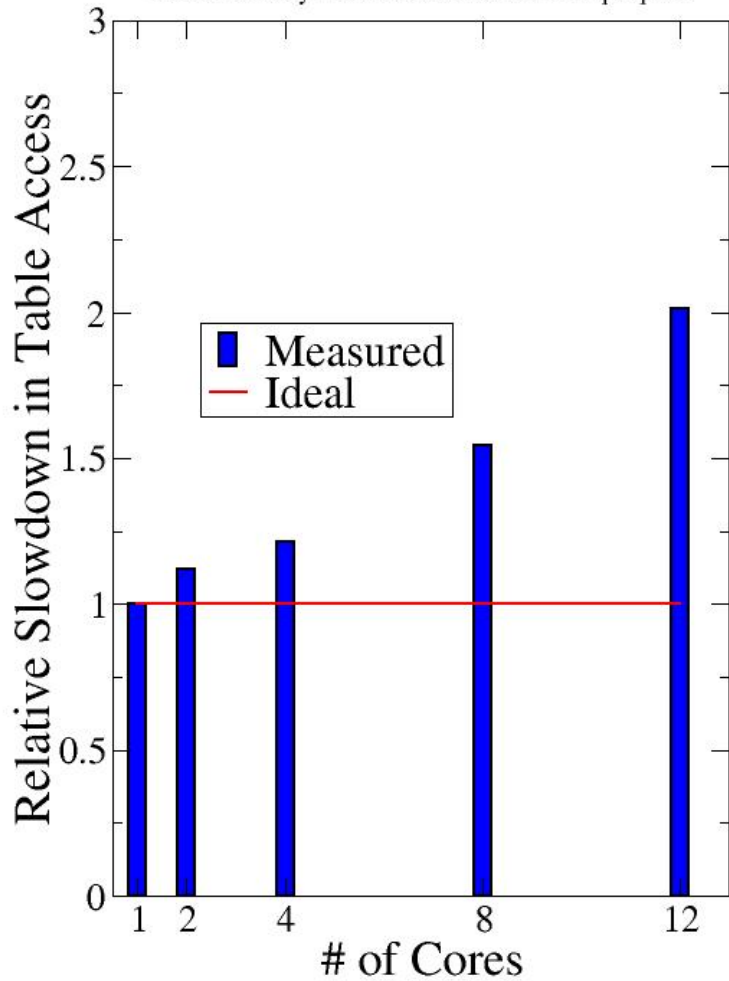


Shen EOS Tabular Access  
Shared Memory -- 2.5 microseconds of work per point



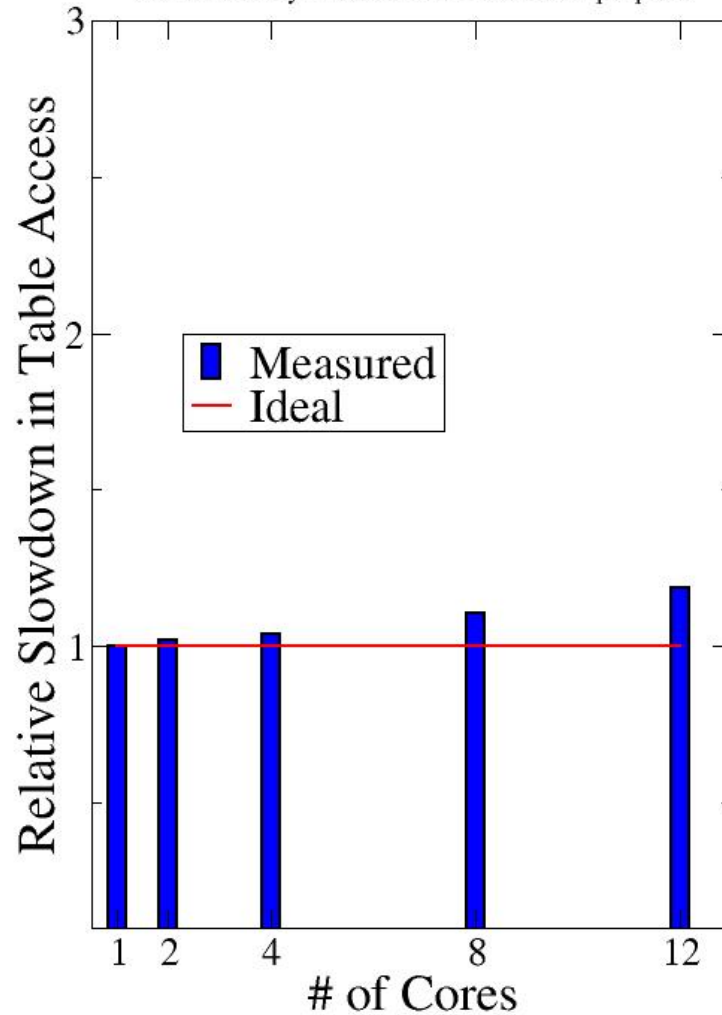
### Shen EOS Tabular Access

Shared Memory -- 2.5 microseconds of work per point



### Shen EOS Tabular Access

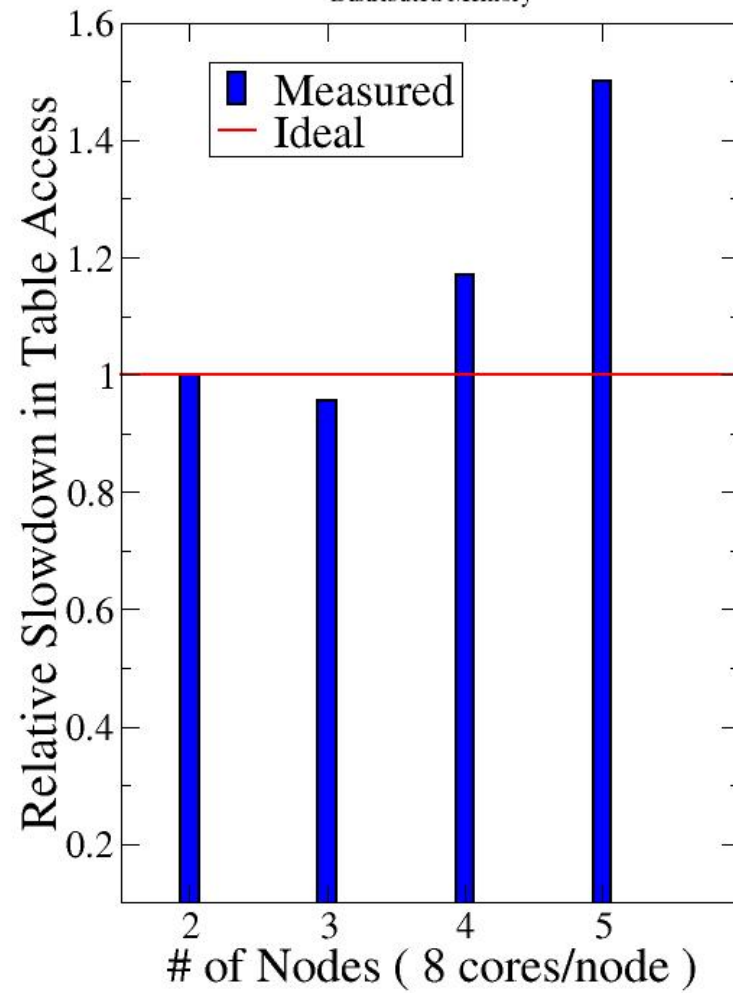
Shared Memory -- 14 microseconds of work per point



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## Shen EOS Tabular Access Distributed Memory

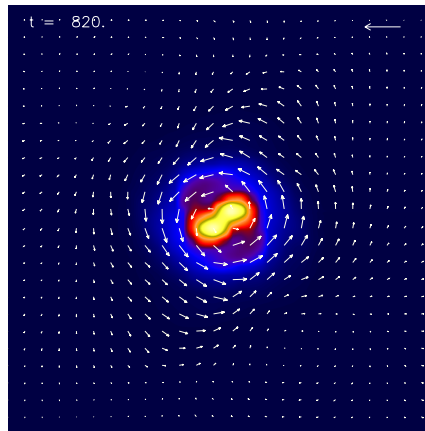


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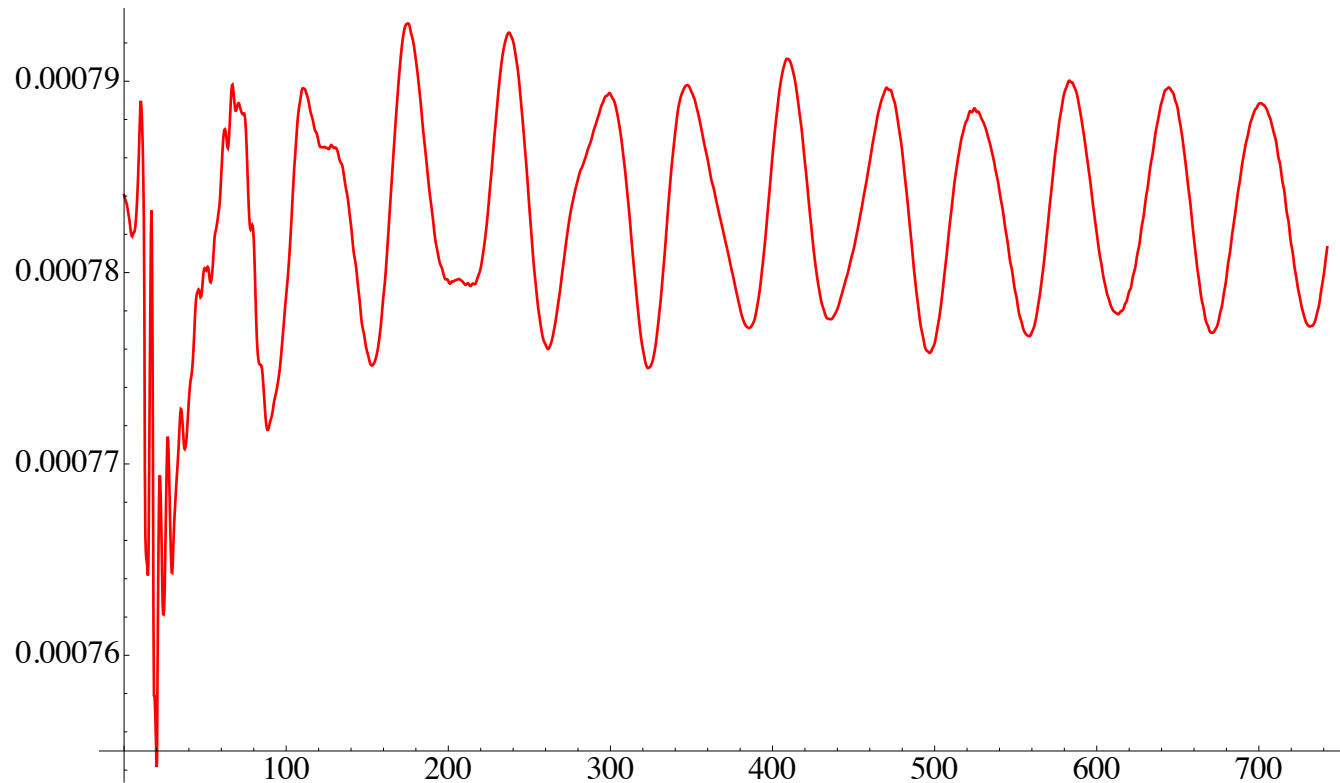
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# MHDe Code Tests

- All C++
- PPM Reconstruction with HLLC Numerical Flux
- AMR
- Equations described in <http://arxiv.org/abs/gr-qc/0605102>



# Preliminary Radial Pulsation Frequencies



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# Conclusions

- Big look-up tables are coming or are already here (EOS, opacity, etc)
- Exaflops computing is coming
- Modern execution models provide tools key for improving scaling constrained codes
- Stay tuned for more

